

Reactor Experiment for Neutrinos and Exotics

Wonjun Lee

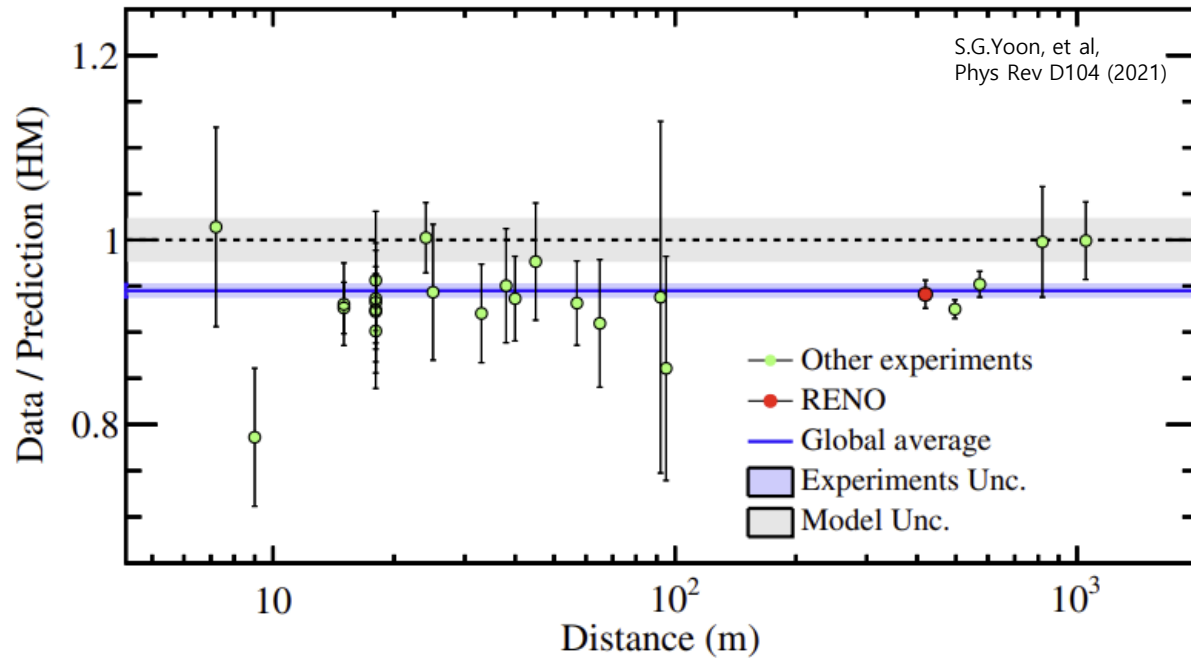
Seoul National University

Reactor Experiment for Neutrinos and Exotics

1. INTRODUCTION

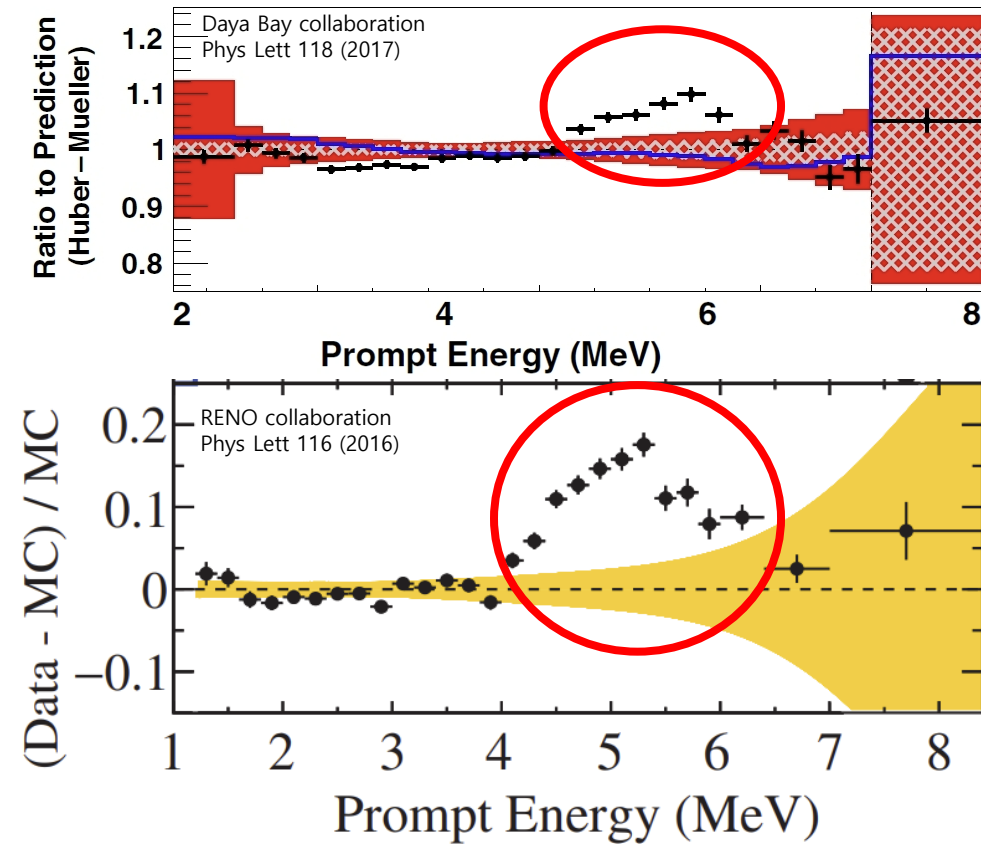
RAA: Reactor Antineutrino Anomaly

Deficit in flux about 5%



5MeV excess

An excess in the prompt spectrum



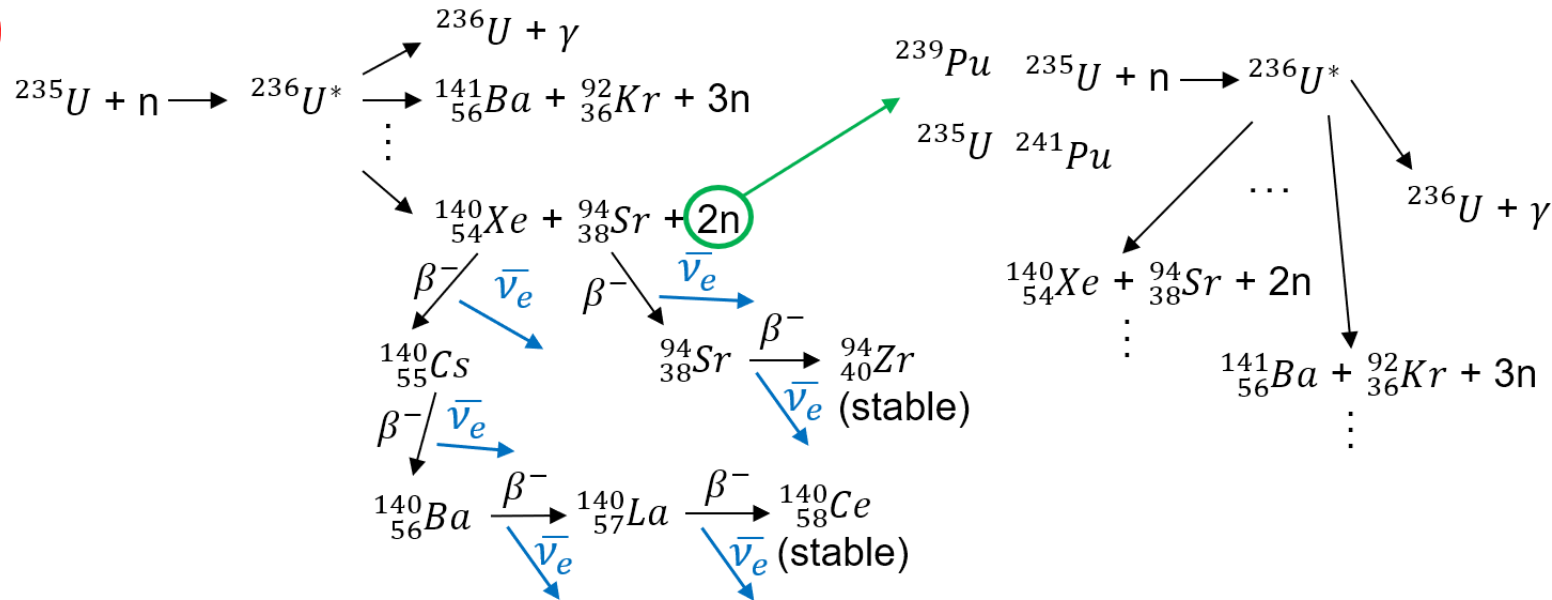
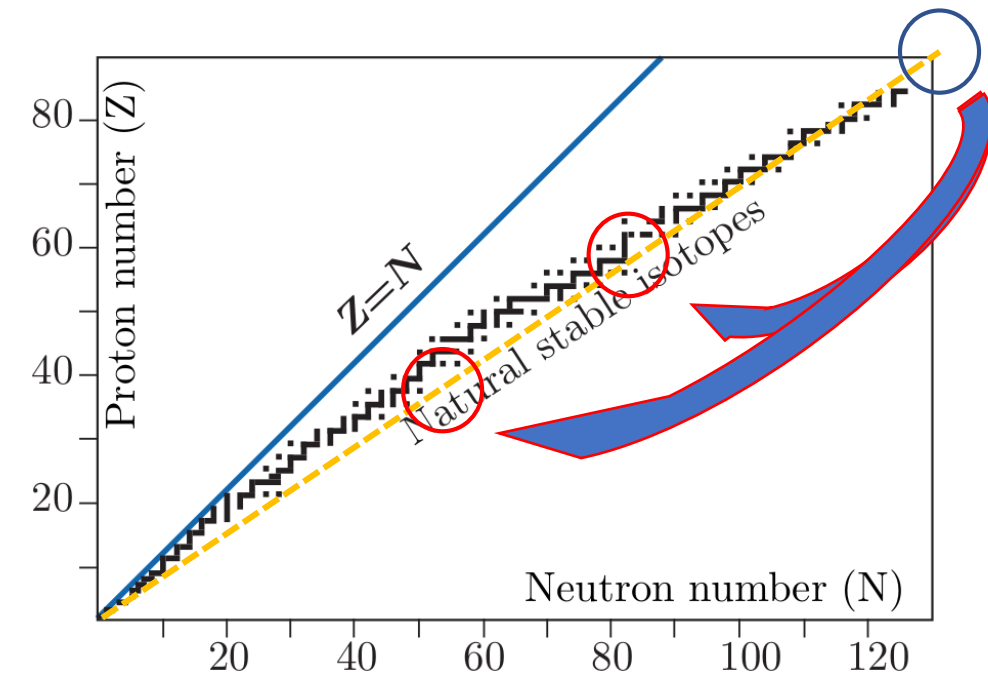
Reactor Experiment for Neutrinos and Exotics

2. REACTOR NEUTRINO

Neutrinos from reactors

By nuclear fission! – Neutron absorption: $\textcircled{n} + \frac{A-1}{Z}X \rightarrow \frac{A}{Z}X^* \rightarrow \text{fission}$

Unstable nuclei go through the beta decay to be stable. $\Rightarrow \nu$



Reactor Experiment for Neutrinos and Exotics

2. REACTOR NEUTRINO

Summation method model

Based on Nuclear databases
(still imperfect...)

$$S_k(E) = \sum_f A_f S_f(E)$$

$$S_f(E) = \sum_b BR_f^b S_f^b$$

$$S_f^b(E_\nu)$$

$e \rightarrow \nu$

Sum

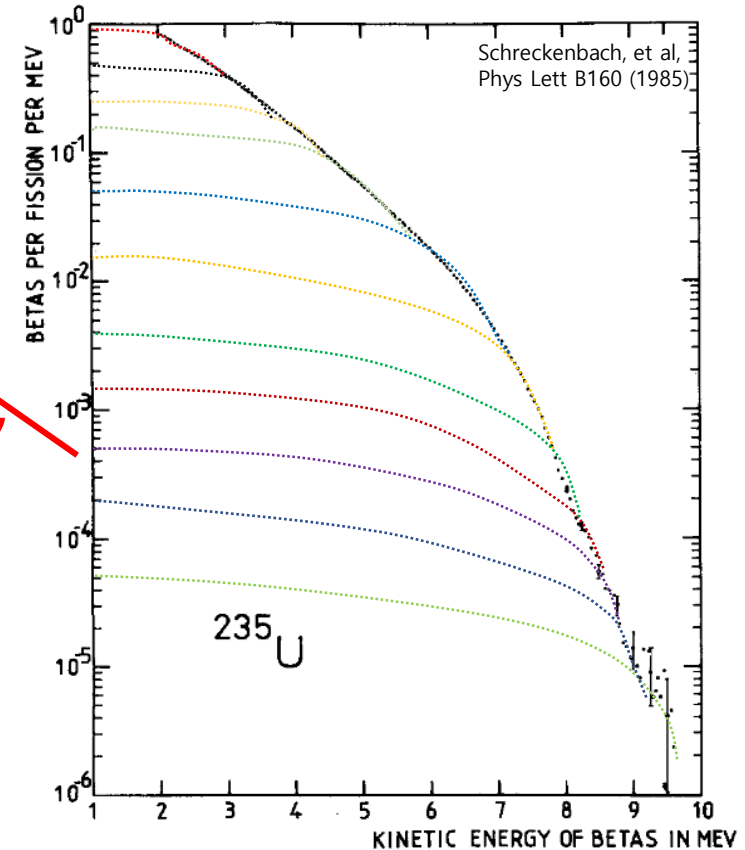
$$S_k(E_\nu)$$

Sum

$$S_i^{\text{vir}}(E_\nu)$$

$e \rightarrow \nu$

Conversion method model



Reactor Experiment for Neutrinos and Exotics

2. REACTOR NEUTRINO

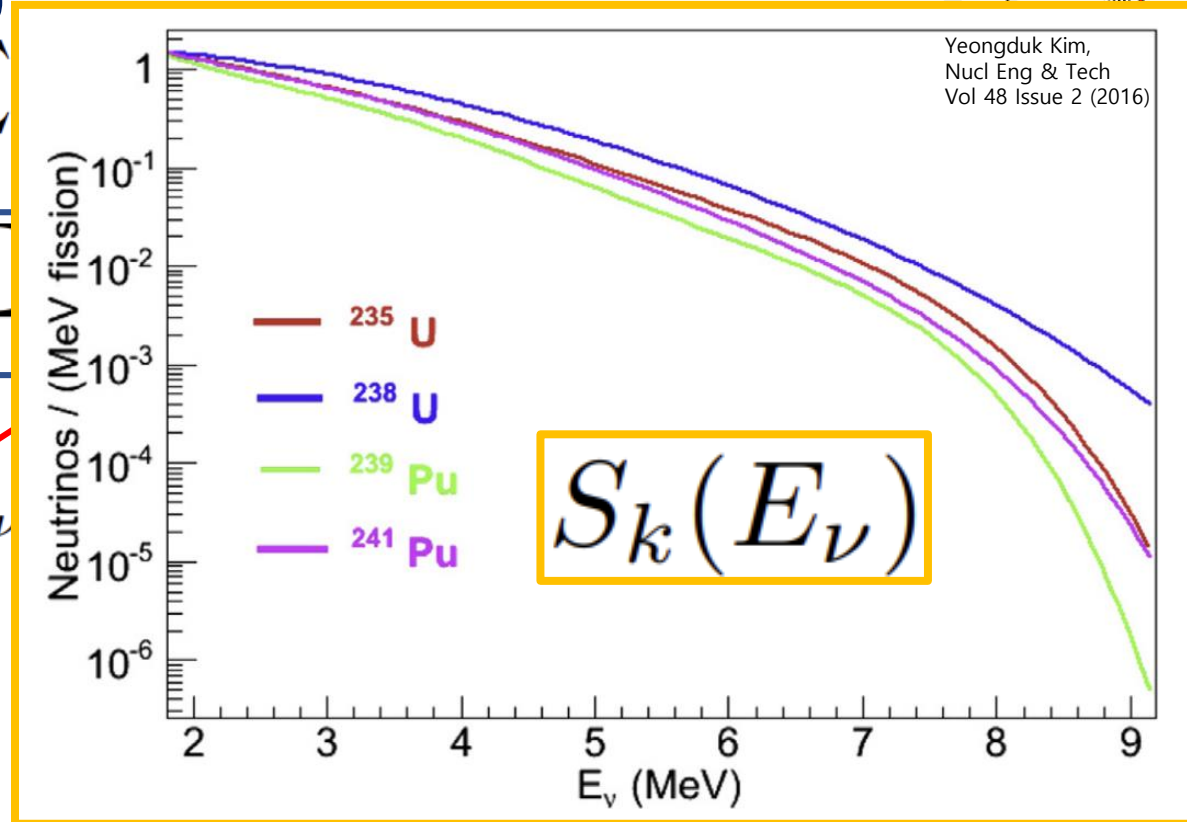
Summation method model

Based on Nuclear databases
(still imperfect....)

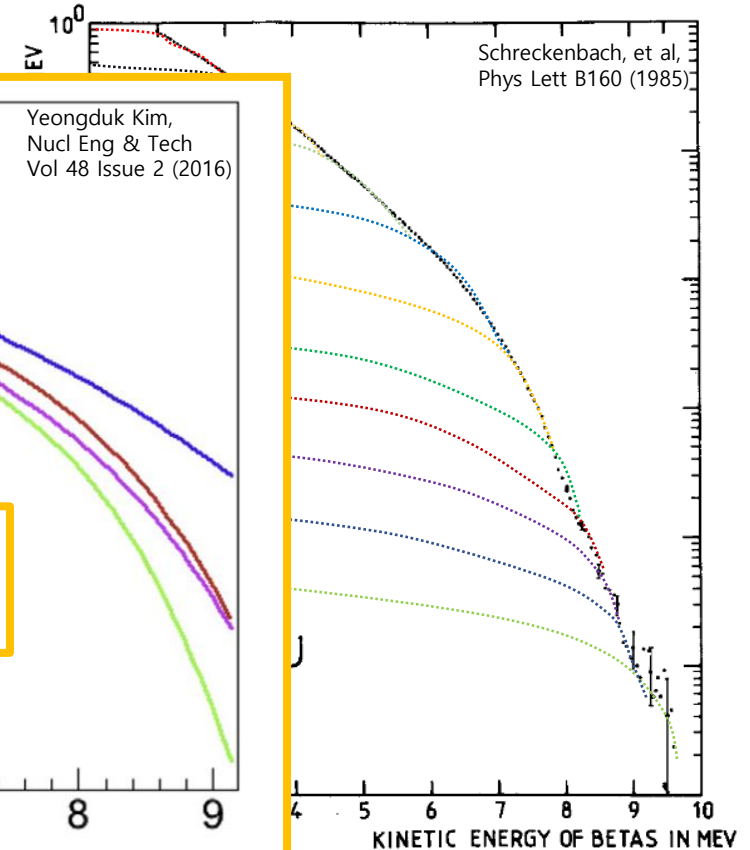
$$S_k(E) = \sum_f$$

$$S_f(E) = \sum_b$$

$$S_f^b(E_\nu)$$



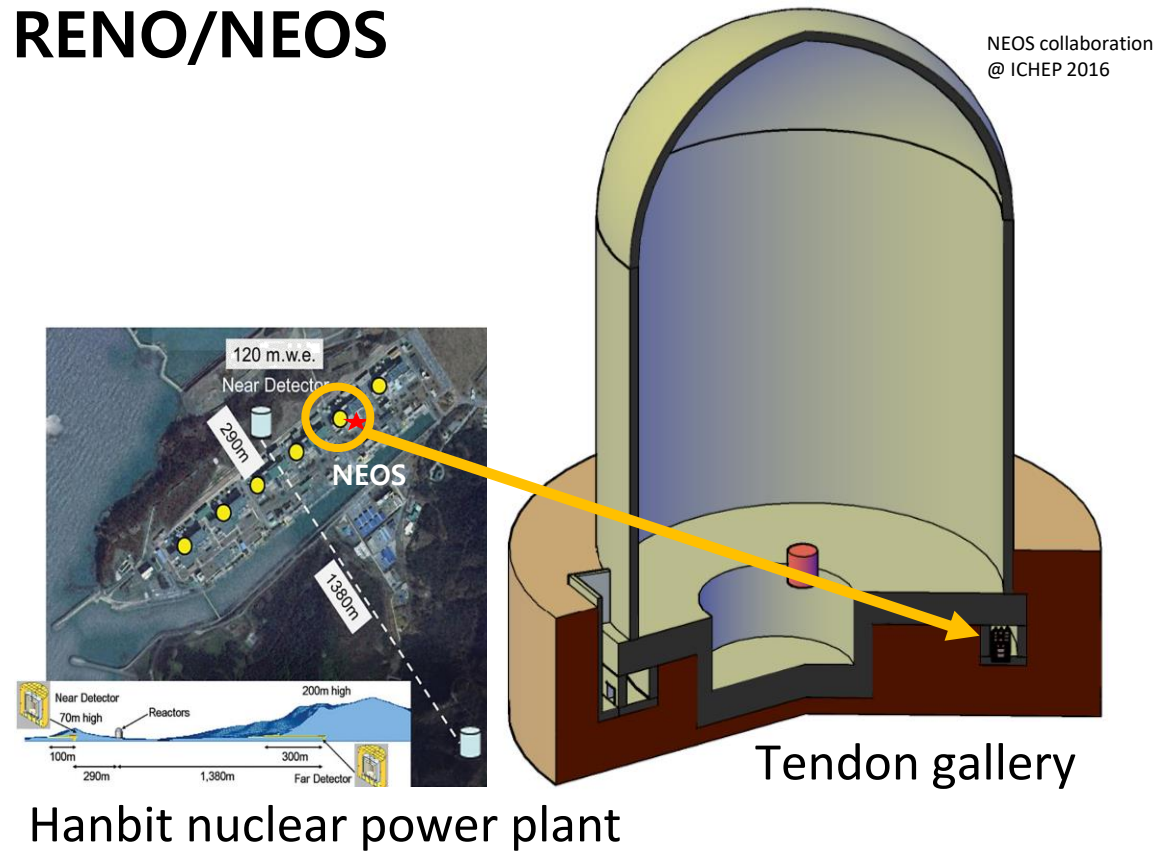
Conversion method model



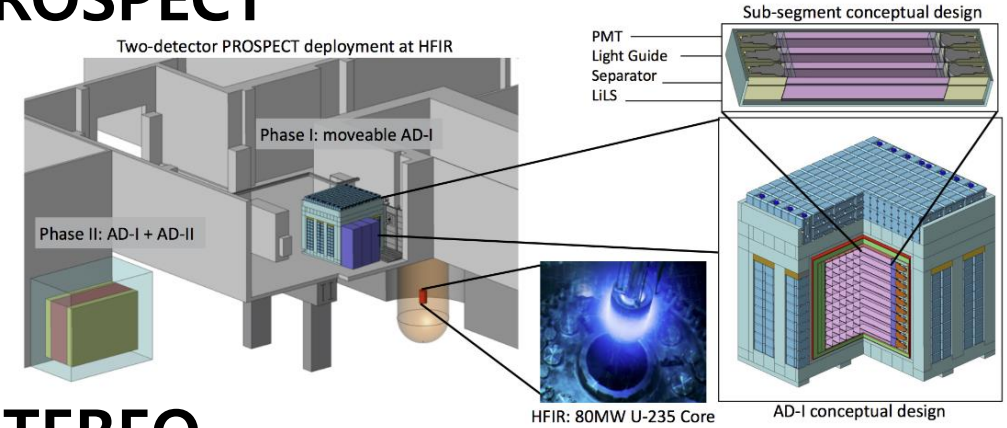
Reactor Experiment for Neutrinos and Exotics

3. REACTOR NEUTRION EXPERIMENTS

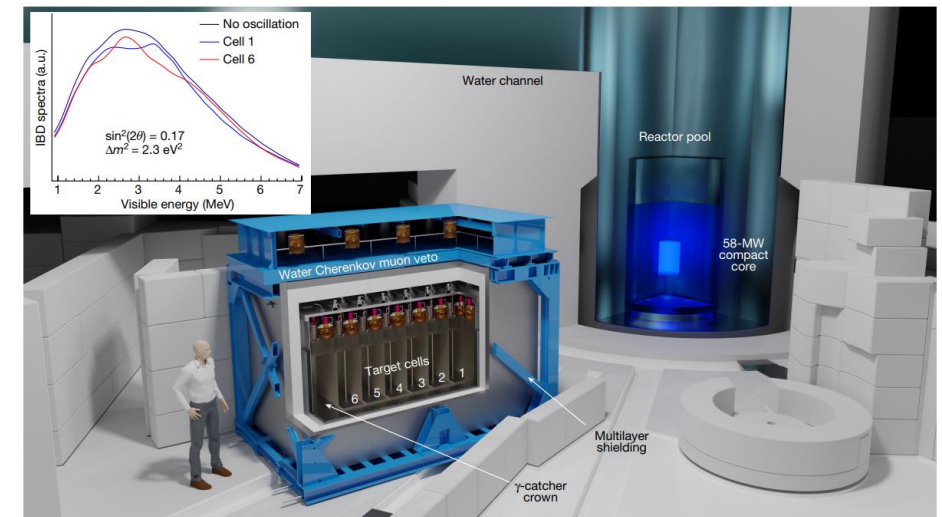
RENO/NEOS



PROSPECT



STEREO

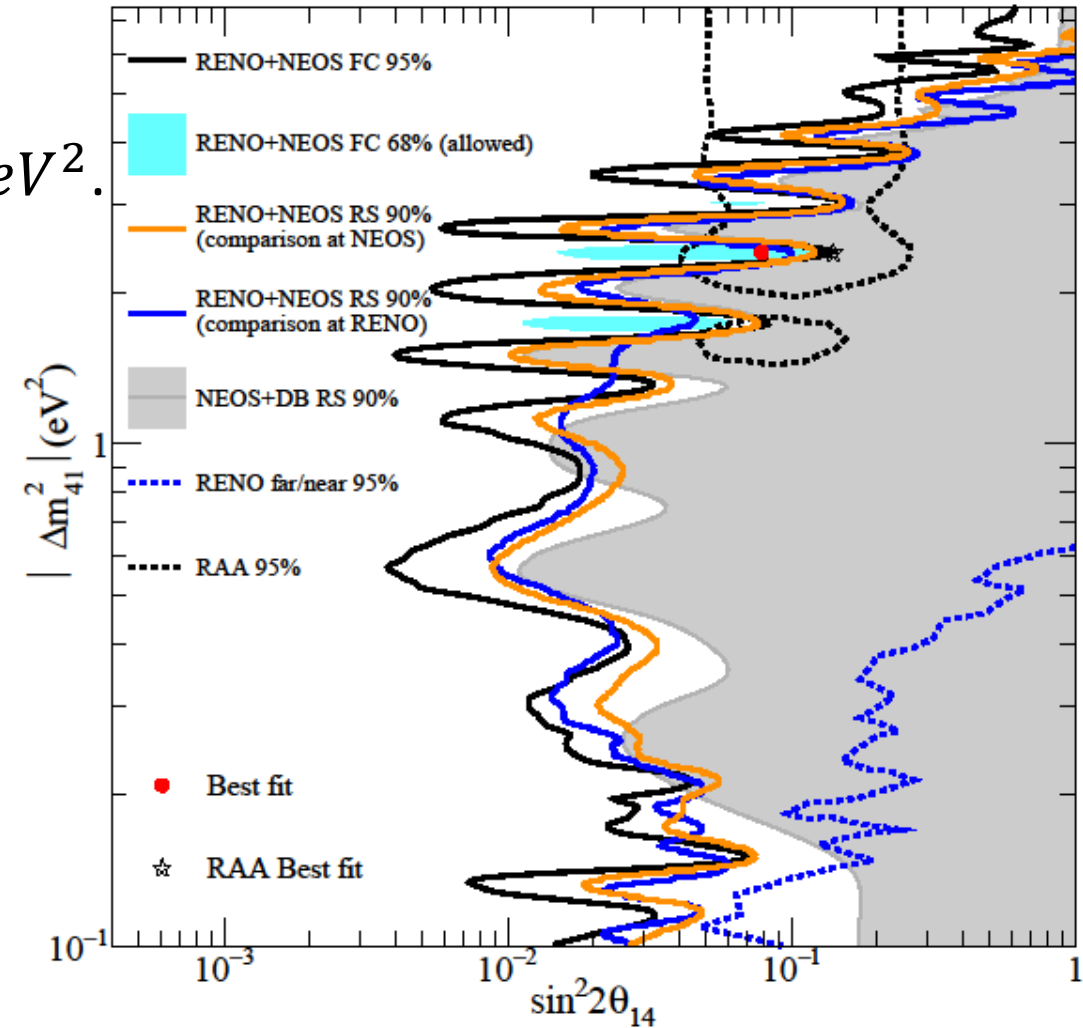
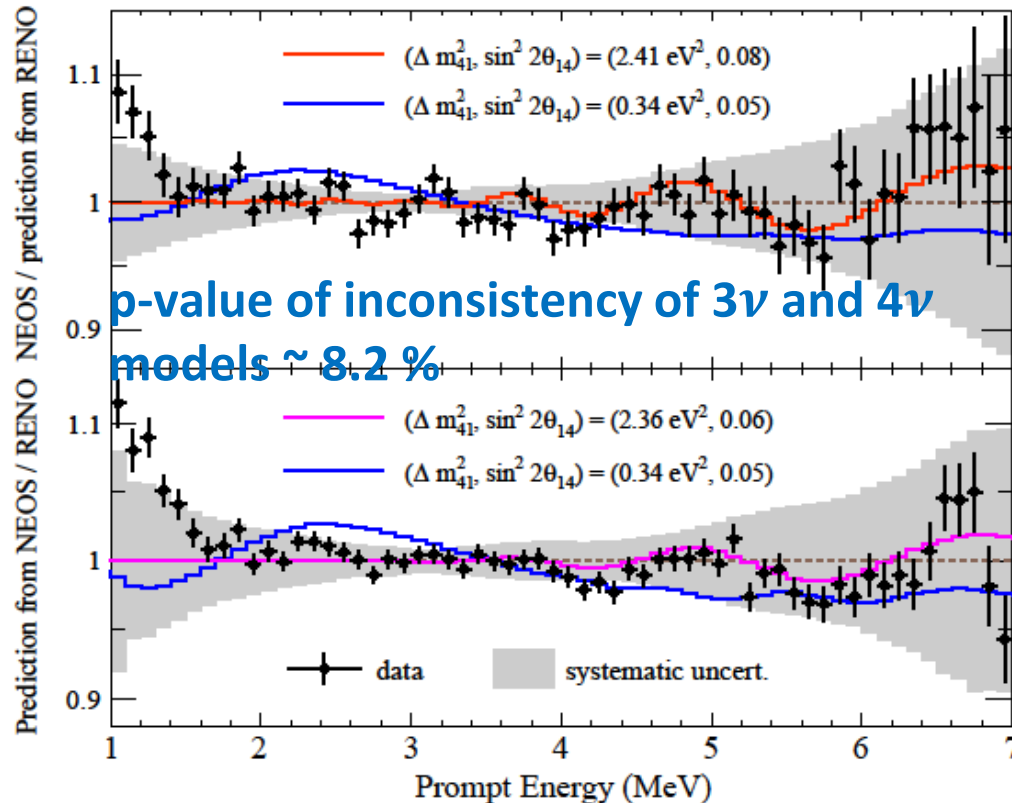


3. REACTOR NEUTRION EXPERIMENTS

RENO/NEOS - LEU

Joint analysis result (PRD 105, 111101)

Hint for the sterile neutrino at $\Delta m_{41}^2 \sim 2 \text{ eV}^2$.

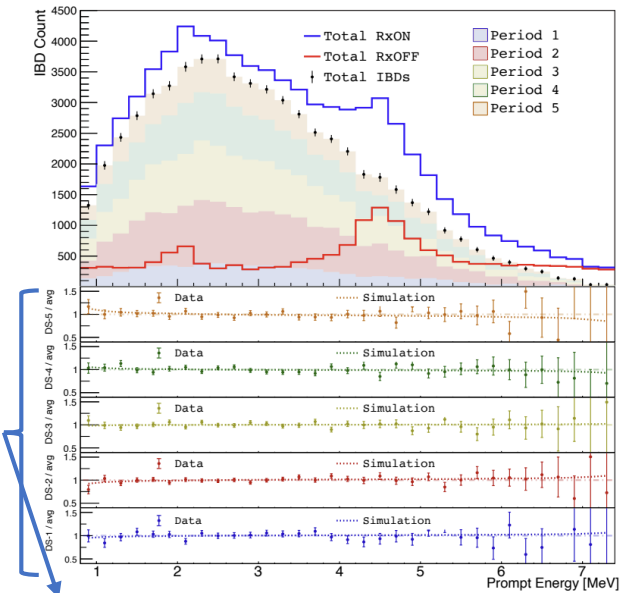


Reactor Experiment for Neutrinos and Exotics

3. REACTOR NEUTRION EXPERIMENTS

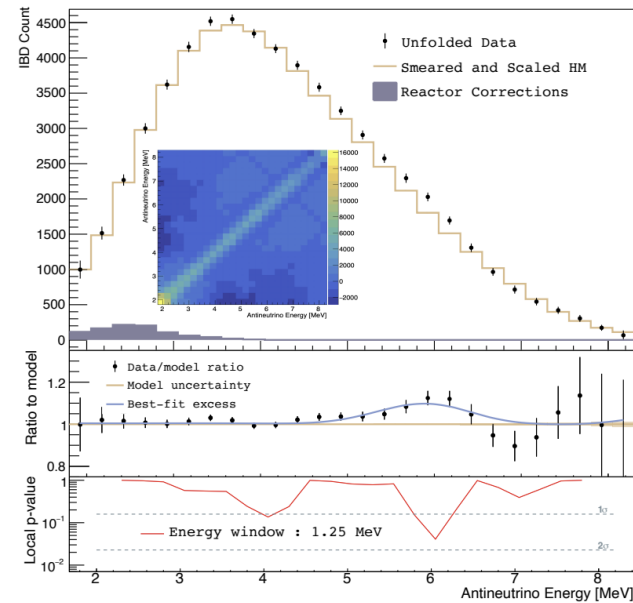
PROSPECT - HEU

Prompt spectrum

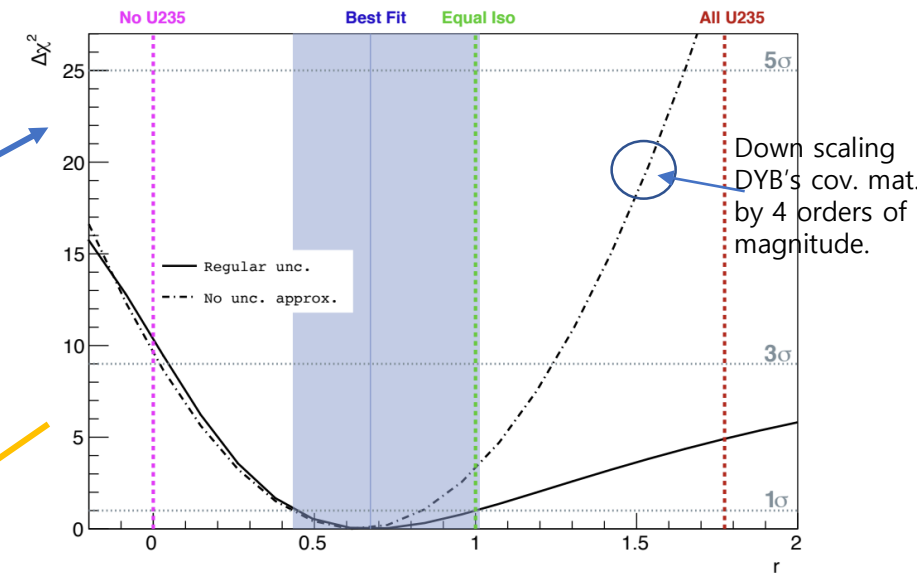


Data(1)/Data(sum)
vs MC(1)/MC(sum)

Neutrino spectrum



1. Gaussian fit to DYB 5MeV excess.
2. Using the fit function, fit PROSPECT 5MeV excess only modifying the amplitude.
3. Get the amplitude ratio "r".



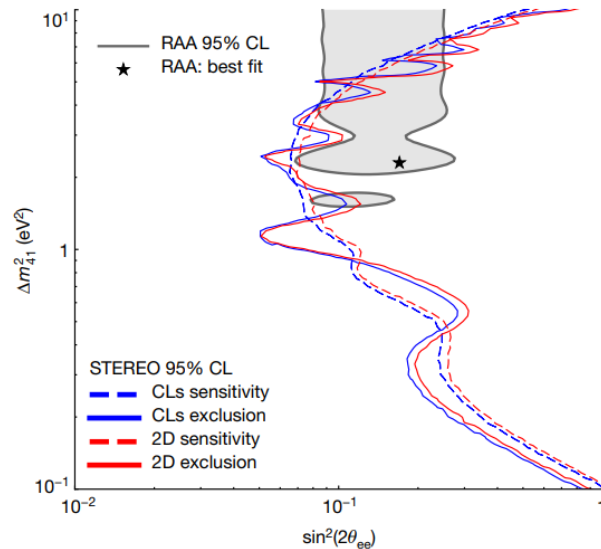
1. Reject the hypothesis that **5MeV excess doesn't come from ^{235}U with 3.2σ .**
2. Reject the hypothesis that **5MeV excess only comes from ^{235}U with 2.2σ .**

3. REACTOR NEUTRION EXPERIMENTS

STEREO - HEU

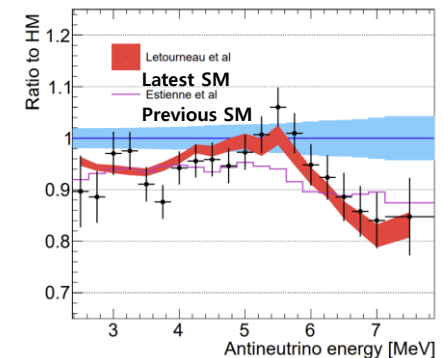
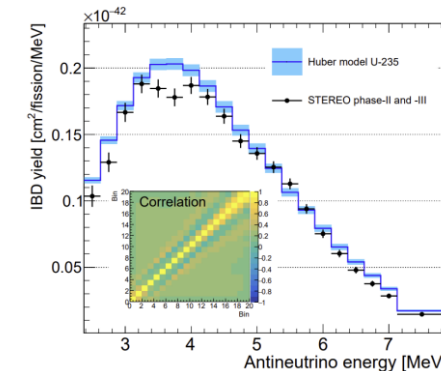
- STEREO data **reject** the “**light sterile neutrino hypothesis**”.

H_{41} 's p-value < 0.05 : rejected
 H_{00} 's p-value = 0.52 : not rejected



- Their data suggest that **the anomalies come from biases in the nuclear experimental data** which are used for the predictions.

For **^{235}U** , **5.5 ± 2.1 % deficit compared to the HM.**
The data **agrees with the recent SM.**



3. REACTOR NEUTRION EXPERIMENTS

Summary of the experiments

RENO/NEOS

- RENO/NEOS gives a hint for the sterile neutrino at $\Delta m_{41}^2 \sim 2 \text{ eV}^2$.

PROSPECT

- PROSPECT concludes that 5MeV excess doesn't come from ^{235}U with 3.2σ and
- 5MeV excess only comes from ^{235}U with 2.2σ .

STEREO

- STEREO data **reject** the “light sterile neutrino hypothesis”.
- Their data suggest that **the anomalies come from biases in the nuclear experimental data** which are used for the predictions.

MOTIVATION

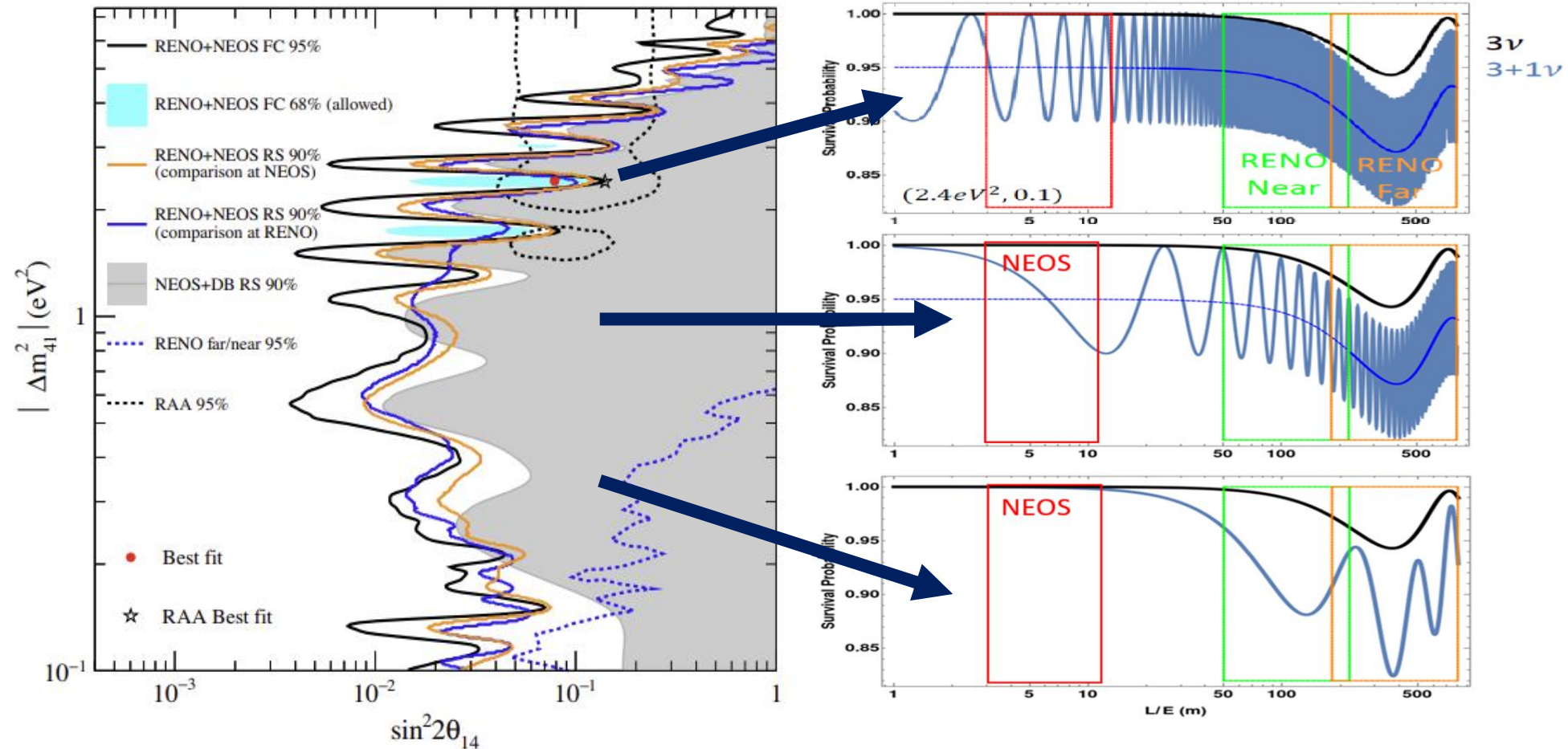
1. Search for the sterile neutrino at $\Delta m_{41}^2 \sim 2 \text{ eV}^2$.
2. Precise measurement for the reactor neutrino spectrum.

Reactor Experiment for Neutrinos and Exotics

4. *RENE* - Reactor Experiment for Neutrinos and Exotics

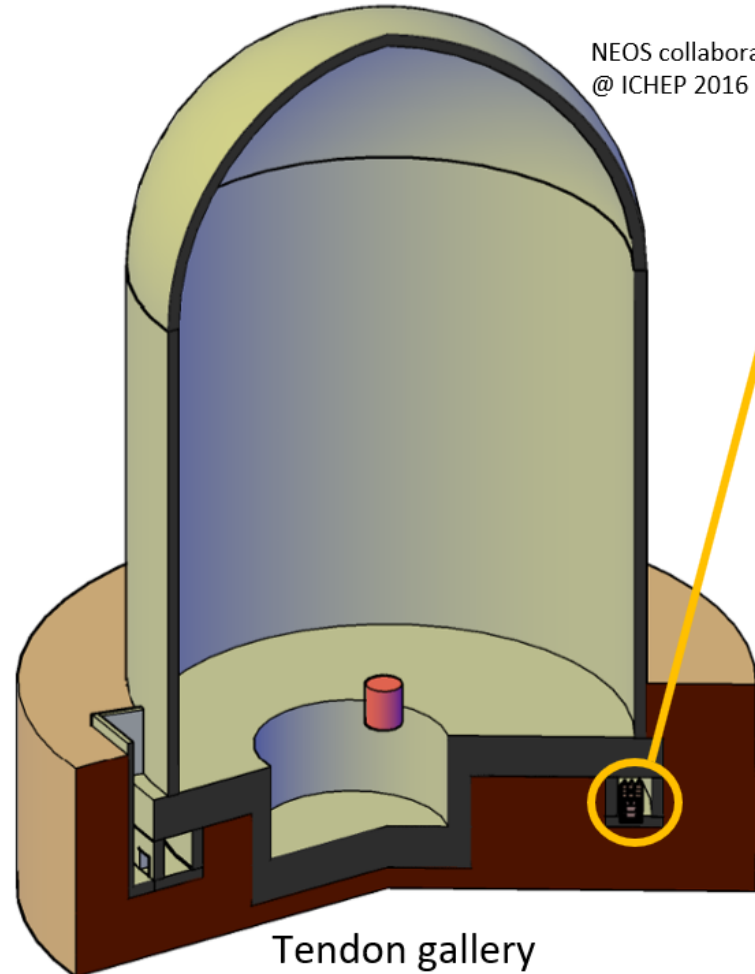
$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 1 - 4 \sum_{i>j} |U_{ei}|^2 |U_{ej}|^2 \sin^2 \left(\frac{\Delta m_{ij}^2 L}{4E} \right) \quad \Delta_{ij} = 1.27 \Delta m_{ij}^2 L/E \text{ (eV}^2 \text{ m/MeV)}$$

$$\simeq 1 - \sin^2 2\theta_{14} \sin^2 \Delta_{41} - \sin^2 2\theta_{13} \sin^2 \Delta_{31},$$



Reactor Experiment for Neutrinos and Exotics

4. *RENE* - Reactor Experiment for Neutrinos and Exotics



NEOS collaboration
@ ICHEP 2016

We are considering **multiple detectors**.

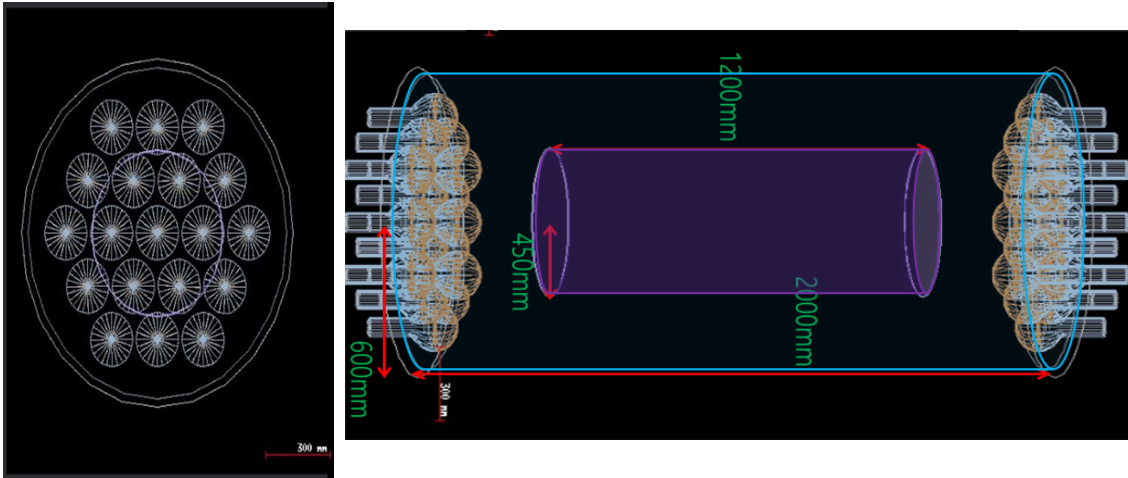
$$\left[\begin{array}{l} \Delta_{ij} = 1.27 \Delta m_{ij}^2 L/E \text{ (eV}^2 \text{ m/MeV)} \\ \pi/1.27 \sim 2.5 \text{ m at } \Delta m^2 \sim 2 \text{ eV}^2 \text{ and } E = 4 \text{ MeV} \end{array} \right]$$

	Baseline (m)
L0	23.6
L1	22.8
L2	21.9
L3	21.7

There are **four candidate locations**, L0, L1, L2, and L3.

Reactor Experiment for Neutrinos and Exotics

4. *RENE* - Reactor Experiment for Neutrinos and Exotics



- Total **38 8-inch pmts** on **Left & Right**
- **450 mm radius** & 0.5% Gd-loaded Target
- => **763.407 L**
- 150 mm thickness Gamma catcher
- **Reflective material** on the wall of the gamma catcher

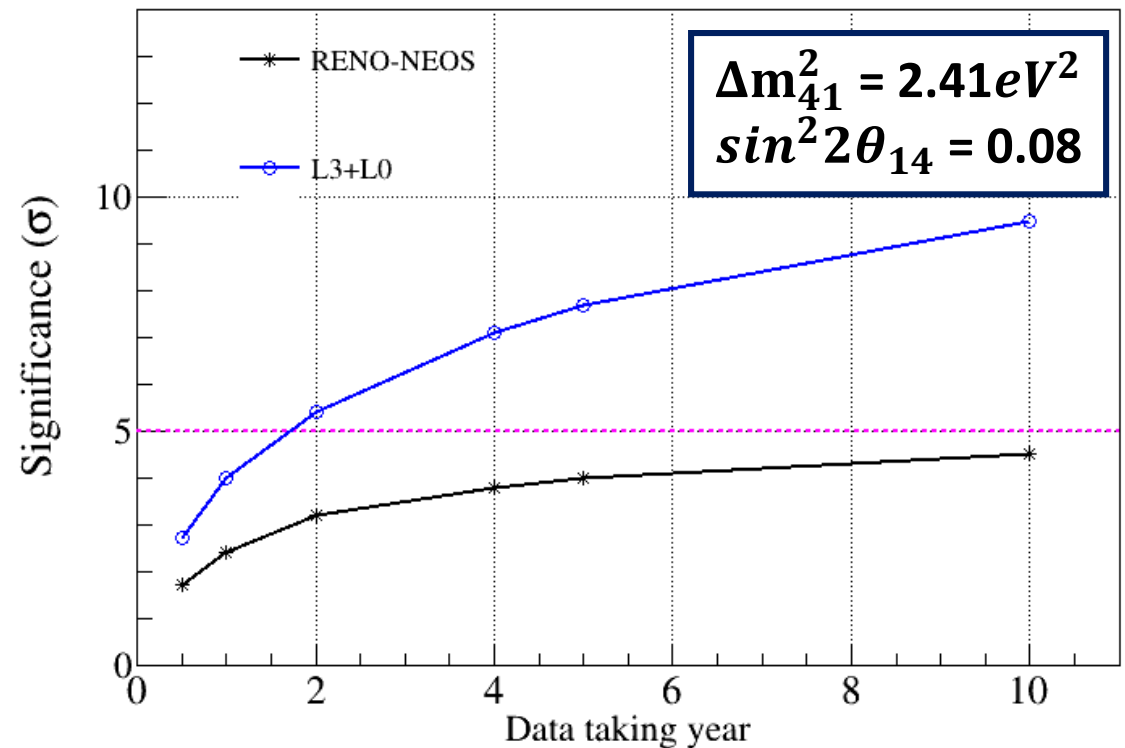
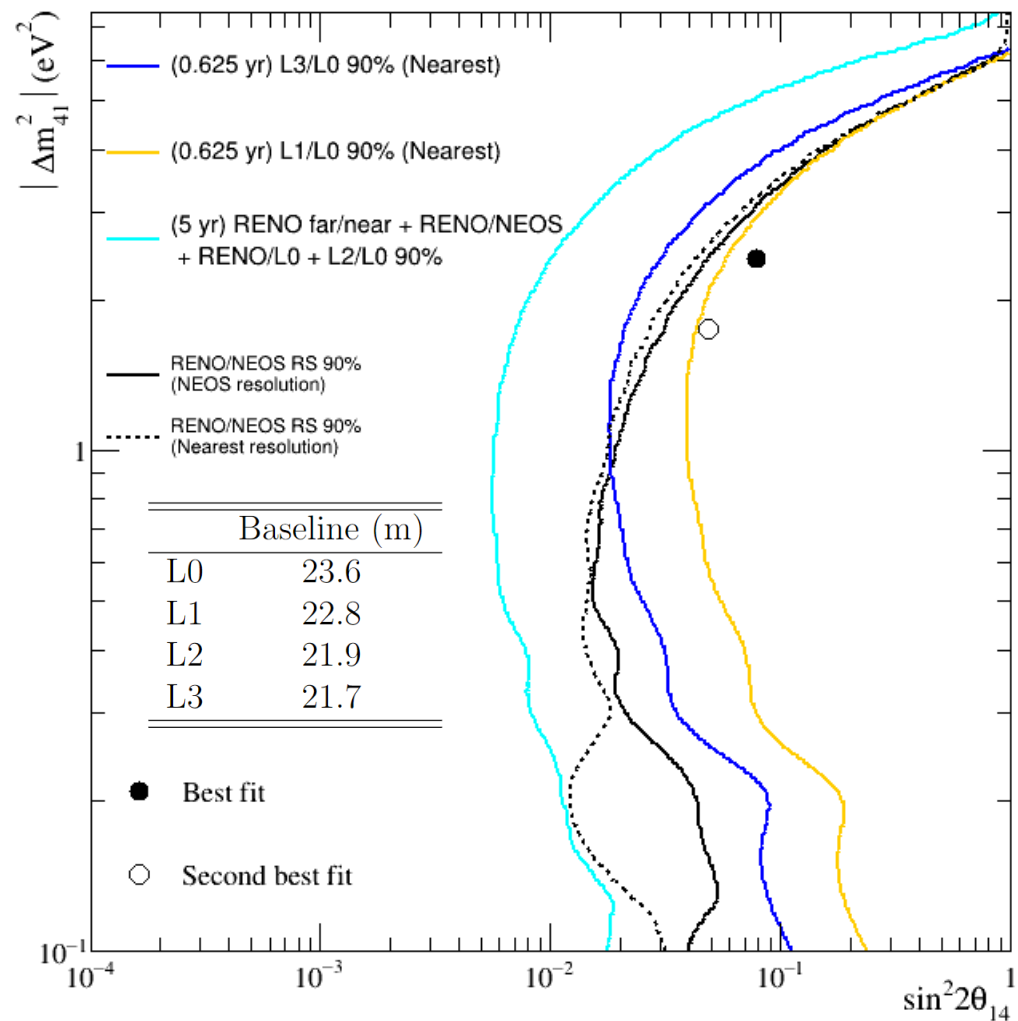
PROTOTYPE



- Total **2 20-inch pmts** on **Left & Right**
- **254 mm radius** & 0.5% Gd-loaded Target
- => **243.220 L**
- 150 mm thickness Gamma catcher
- **Reflective material** on the wall of the gamma catcher

Reactor Experiment for Neutrinos and Exotics

4. *RENE* - Reactor Experiment for Neutrinos and Exotics



4. *RENE* - Reactor Experiment for Neutrinos and Exotics

- RENO-NEOS joint analysis shows an interesting hint **for the existence of the sterile neutrino at $\Delta m_{41}^2 \sim 2 \text{ eV}^2$.**
- To search the mass parameter region, we are considering **multiple detectors** at nearest sites in the Hanbit nuclear power plant.
- **We expect to search for the sterile neutrino at the above parameter region.**
- **After that, we have a future plan to precisely measure the neutrino spectrum at LEU and HEU reactors with the same detector.**

**감사합니다.
Thank you.**

BACKUP

Reactor Experiment for Neutrinos and Exotics

* *Backup*

Fission: Break up of a nucleus

$${}^A_Z X \rightarrow {}^{a_1}_{z_1} M + {}^{a_2}_{z_2} N + Q_{fission} \quad \text{where } A = a_1 + a_2 \text{ and } Z = z_1 + z_2$$

In the rest frame,

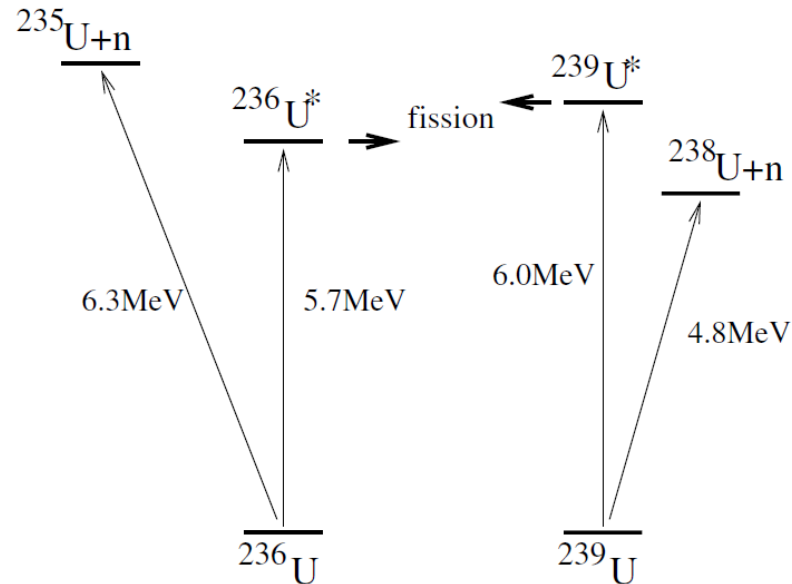
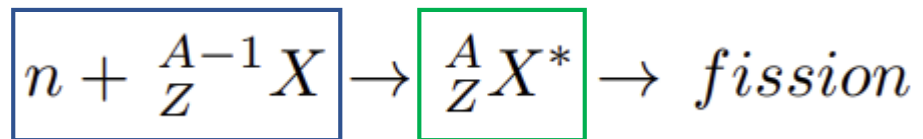
$$m(A, Z) = Zm_p + Nm_n - B(A, Z) \quad \text{BE : The negative difference of the nuclear mass and the masses of the constituents.}$$

$$B(A, Z) = a_v A - a_s A^{2/3} - a_c \frac{Z^2}{A^{1/3}} - a_a \frac{(N-Z)^2}{A} + \delta(A), \quad \delta(A) = \begin{cases} 33.6 A^{-3/4} & \text{if } N \text{ and } Z \text{ are even} \\ -33.6 A^{-3/4} & \text{if } N \text{ and } Z \text{ are odd} \\ 0 & \text{if } A = N + Z \text{ is odd} \end{cases}$$

Reactor Experiment for Neutrinos and Exotics

* Backup

- Odd N fissile material > Even N fissile material



$$Q_{fis} = m_n + T_n(A-1) + Zm_z + (N-1)m_n - B(A-1, Z) - [Zm_z + Nm_n - B(A, Z) + E_{ex}] = T_n(A-1) - E_{ex} + B(A, Z) - B(A-1, Z),$$

$$T_n(A-1) = Q_{fis} + E_{ex} - S_n(A) = \Delta E(A) - S_n(A),$$

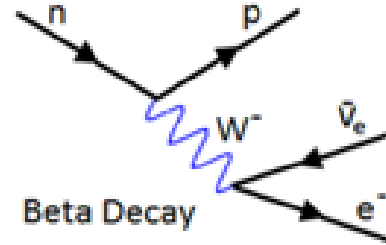
$$\delta(A) = \begin{cases} 33.6A^{-3/4} & \text{if } N \text{ and } Z \text{ are even} \\ -33.6A^{-3/4} & \text{if } N \text{ and } Z \text{ are odd} \\ 0 & \text{if } A = N + Z \text{ is odd} \end{cases}$$

$$\delta(A) - \delta(A-1)$$

Reactor Experiment for Neutrinos and Exotics

* Backup

β decay



Start from Fermi's Golden rule

$$\lambda = \frac{2\pi}{\hbar} |V_{fi}|^2 \rho(E_f)$$

$$V_{fi} = g \int \underbrace{[\psi_f^* \psi_e^* \psi_\nu^*]}_{\text{final state}} O_x \psi_i dv$$

The allowed approximation

$$\psi_e = \frac{e^{i\mathbf{p}\cdot\mathbf{r}/\hbar}}{\sqrt{V}} = \frac{1 + \frac{i\mathbf{p}\cdot\mathbf{r}}{\hbar} + \dots}{\sqrt{V}} \simeq \frac{1}{\sqrt{V}}$$
$$\psi_\nu = \frac{e^{i\mathbf{q}\cdot\mathbf{r}/\hbar}}{\sqrt{V}} = \frac{1 + \frac{i\mathbf{q}\cdot\mathbf{r}}{\hbar} + \dots}{\sqrt{V}} \simeq \frac{1}{\sqrt{V}}.$$

=> **allowed transition**

If we consider higher orders of $\mathcal{O}(\mathbf{p})$ & $\mathcal{O}(\mathbf{q})$,
=> **forbidden transition**

Reactor Experiment for Neutrinos and Exotics

* Backup

β decay

The partial decay rate

$$d\lambda = \frac{2\pi}{\hbar} |V_{fi}|^2 \frac{d^2 n}{dE_f}$$

$d^2 n = (4\pi)^2 \frac{p^2 dp q^2 dq}{h^6} V^2$

$$= \frac{2\pi}{\hbar} [g \int (\psi_f^* \psi_e^* \psi_\nu^*) O_x \psi_i dv]^2$$

$$\times (4\pi)^2 \frac{p^2 dp q^2 dq}{h^6 dE_f} V^2$$

$$= \frac{2\pi}{\hbar} g^2 \left[\int \psi_f^* O_x \psi_i dv \right]^2 (4\pi)^2 \frac{p^2 dp q^2 dq}{h^6 dE_f}$$

$$= \frac{2\pi}{\hbar} g^2 |M_{fi}|^2 (4\pi)^2 \frac{p^2 q^2}{h^6} \frac{dp dq}{dE_f},$$

$$\psi_e = \frac{e^{i\mathbf{p}\cdot\mathbf{r}/\hbar}}{\sqrt{V}} = \frac{1 + \frac{i\mathbf{p}\cdot\mathbf{r}}{\hbar} + \dots}{\sqrt{V}} \simeq \frac{1}{\sqrt{V}}$$
$$\psi_\nu = \frac{e^{i\mathbf{q}\cdot\mathbf{r}/\hbar}}{\sqrt{V}} = \frac{1 + \frac{i\mathbf{q}\cdot\mathbf{r}}{\hbar} + \dots}{\sqrt{V}} \simeq \frac{1}{\sqrt{V}}.$$

E_f is $E_e + E_\nu = E_e + qc$
 $dq/dE_f = 1/c$ at fixed E_e

Reactor Experiment for Neutrinos and Exotics

* *Backup*

β decay

The # of e in the range p and $p+dp$

$$N_{\beta}(p) dp = A p^2 q^2 dp, \text{ where } A = \frac{2\pi}{h} g^2 |M_{fi}|^2 (4\pi)^2 / (h^6 c)$$

$$N_{\beta}(p) = \frac{A}{c} p^2 (Q - T_e)^2$$

$E_{\nu} = Q - T_e$ the decay energy Q

$$= \frac{A}{c} p^2 (E_0 - E_e)^2$$

$$= \frac{A}{m_e^2 c^5} p^2 (W_0 - W)^2,$$

where W is $\frac{E_e}{m_e c^2} + 1$

$$N_{\beta}(W) = K p^2 (W - W_0)^2 F(Z, W)$$

The influence of the
nuclear Coulomb potential

Reactor Experiment for Neutrinos and Exotics

* *Backup*

β decay

The more correct β spectrum formula

$$N_{\beta}(W) = K p^2 (W - W_0)^2 F(Z, W) \\ \times L_0(Z, W) C(Z, W) S(Z, W) \\ \times G_{\beta}(Z, W) (1 + \delta_{WM} W),$$

The weak interaction
correction for the finite size
of the nucleus

The EM interaction
correction for the finite size
of the nucleus

The radiative correction
for the virtual and real
photon emissions

The weak magnetism
correction

The screening effect of the
nuclear charge by bound
electrons

Radiative correction

It is crucial to correctly account for both the virtual and real photon contributions to ensure the proper cancellation of 3. Almost all decays relevant in our problem are of this type and the corresponding expression for Fermi decays is numerically very close. divergences.

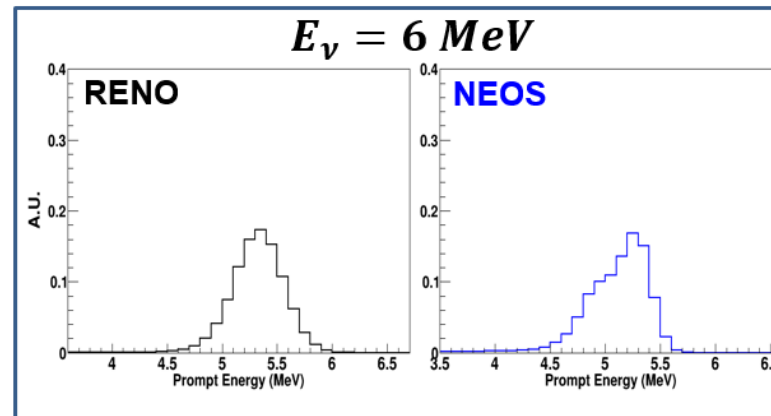
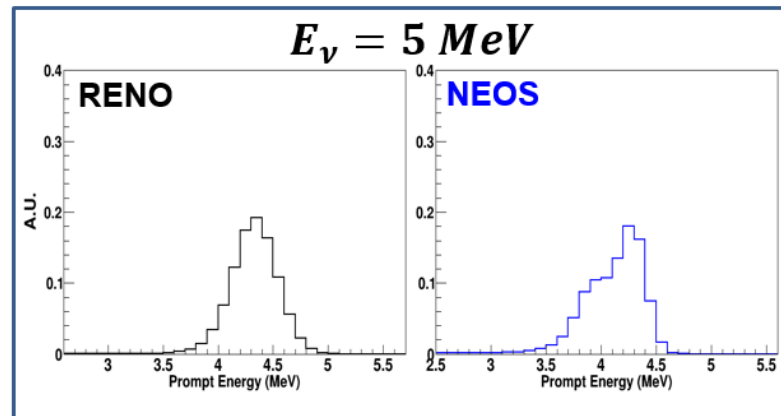
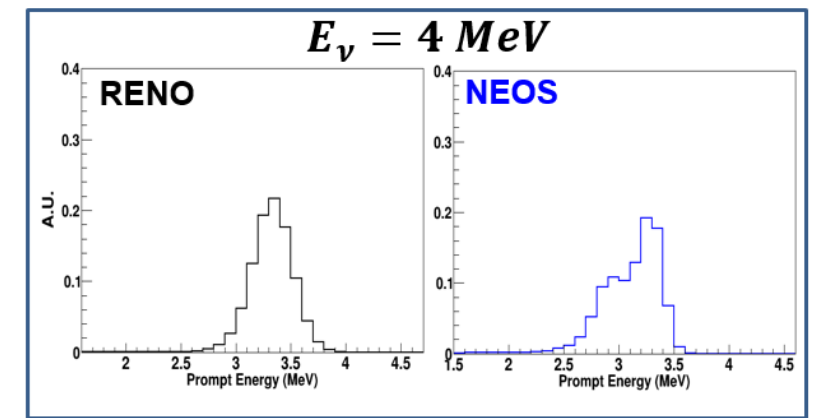
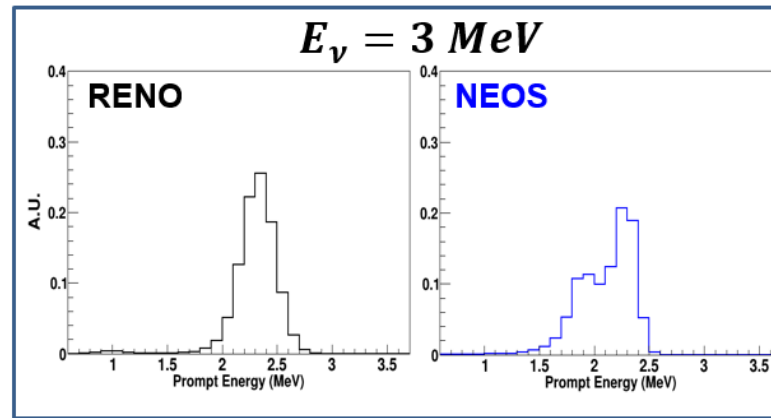
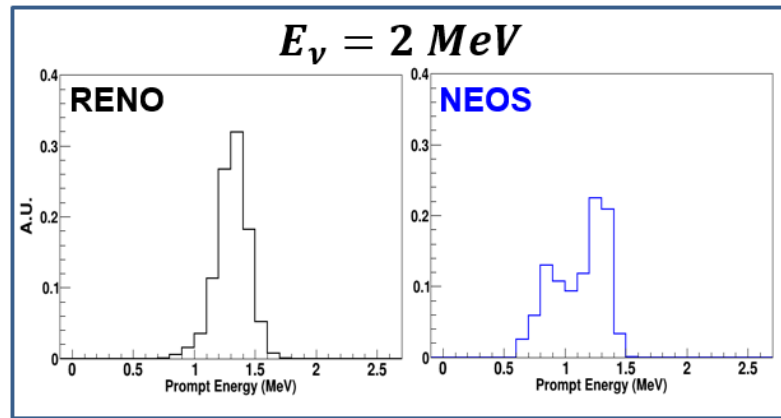
Weak magnetism

Interaction btw current and magnetic moment of nucleon

Reactor Experiment for Neutrinos and Exotics

* *Backup*

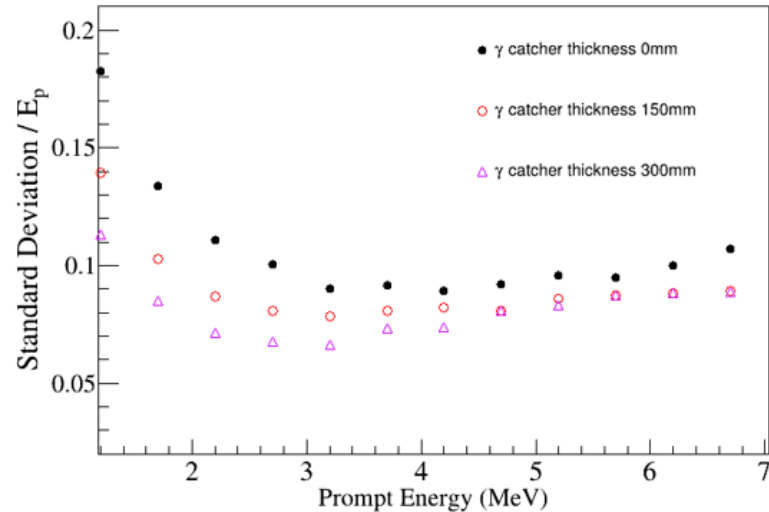
RENO & NEOS detector **responses** of the monochromatic neutrino energy
(PRD 104, 111301) (PRD 105, 111101)



NEOS's **low energy 2nd peak**
is originated by **escaping**
gammas from the target.

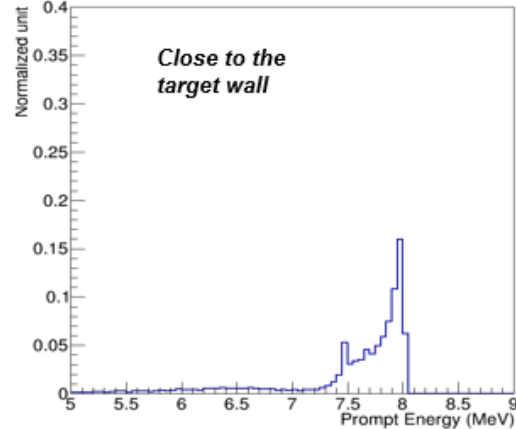
Reactor Experiment for Neutrinos and Exotics

* Backup

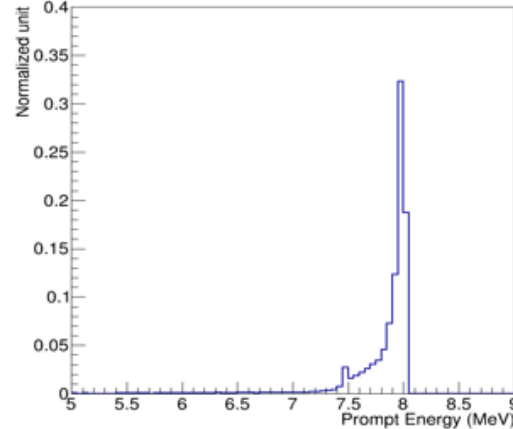


@ 7.5 MeV positron

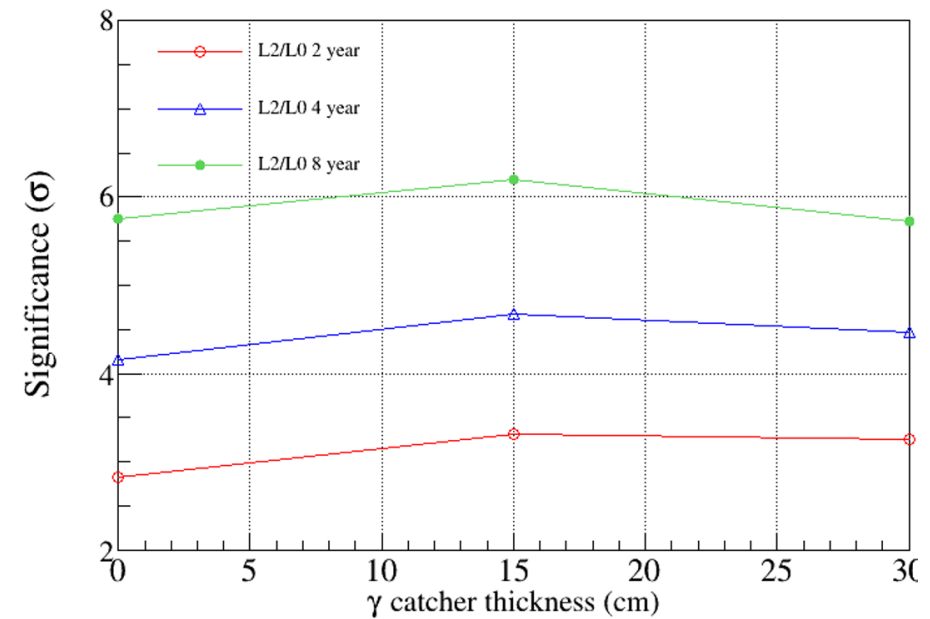
400mm < IBD vertex < 450mm



IBD vertex < 400mm



Under the size constraint and the statistics, 150mm thick gamma catcher case would be the best choice.

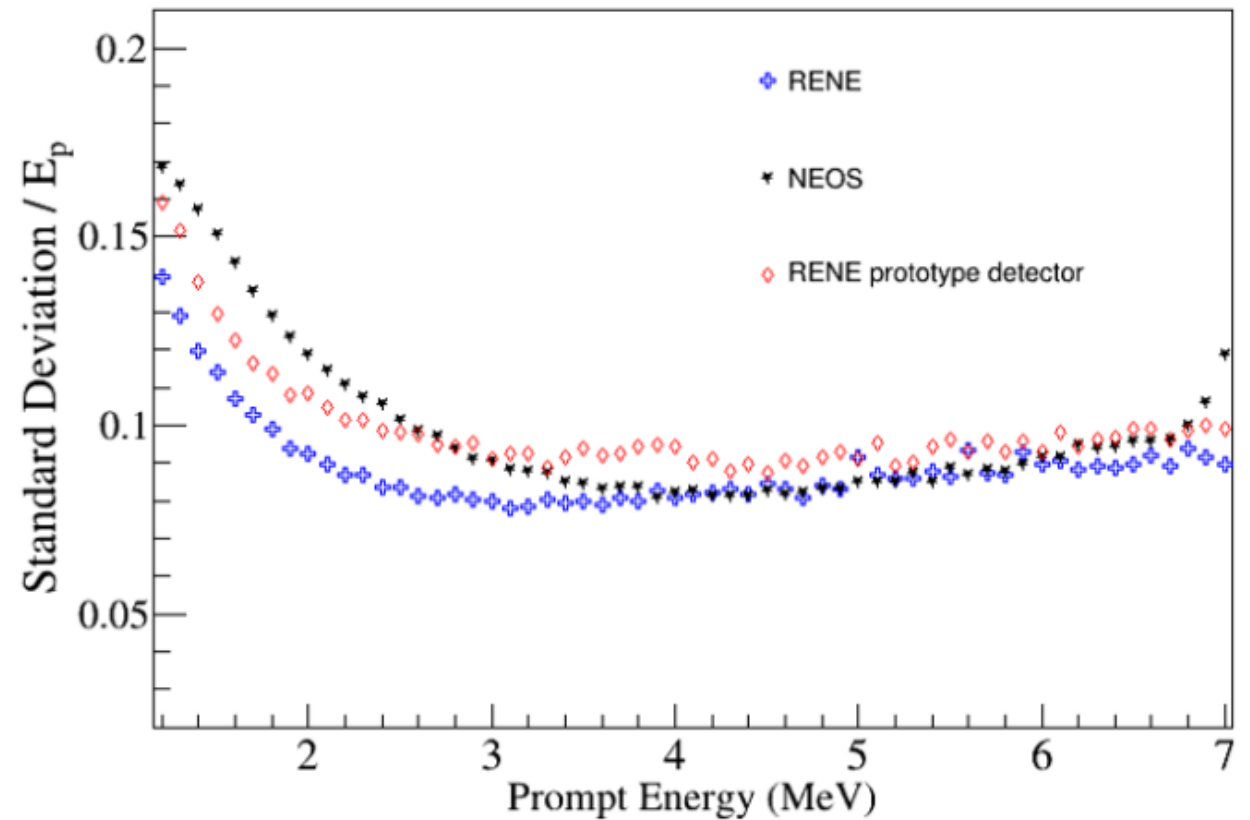


Reactor Experiment for Neutrinos and Exotics

* *Backup*

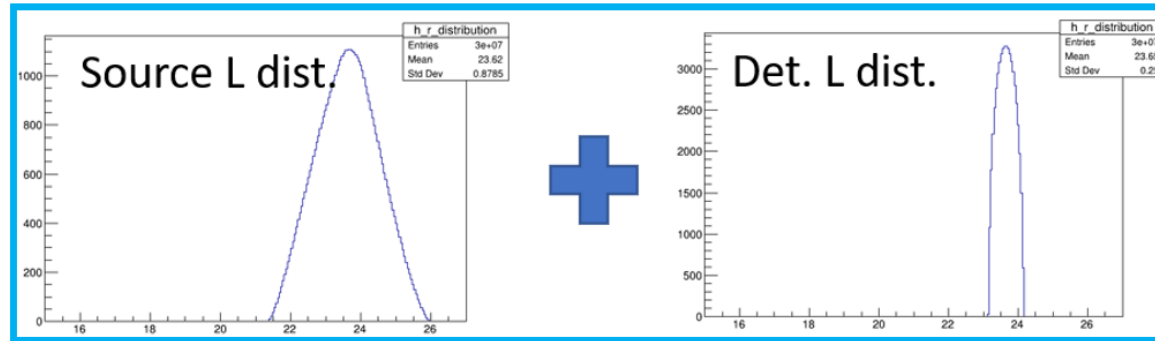


- Total **2 20-inch pmts** on **Left & Right**
- **254 mm radius** & 0.5% Gd-loaded Target
- => **243.220 L**
- 150 mm thickness Gamma catcher
- **Reflective material** on the wall of the gamma catcher

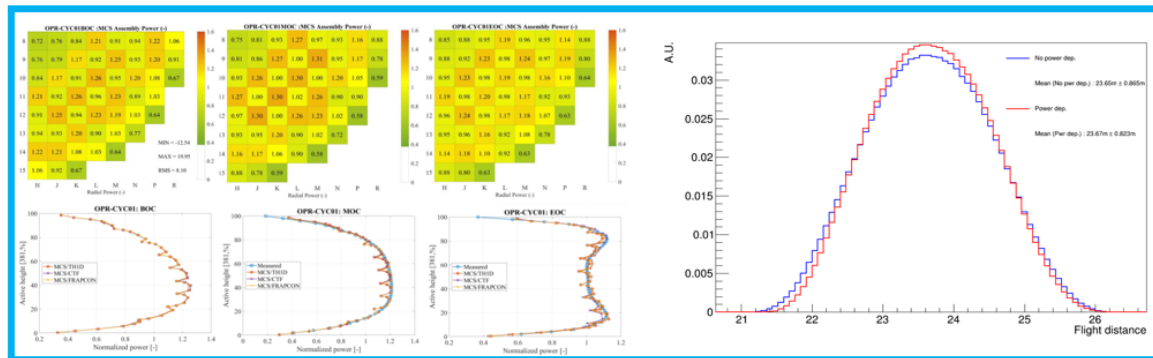
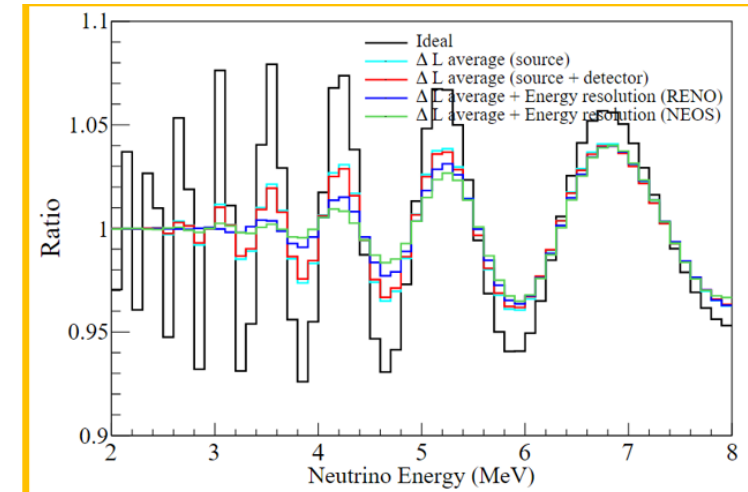


Reactor Experiment for Neutrinos and Exotics

* Backup



Reactor source core dimension effect is larger than the effect of the detector size.



The source core's power distribution doesn't affect to the prompt energy spectrum.

